







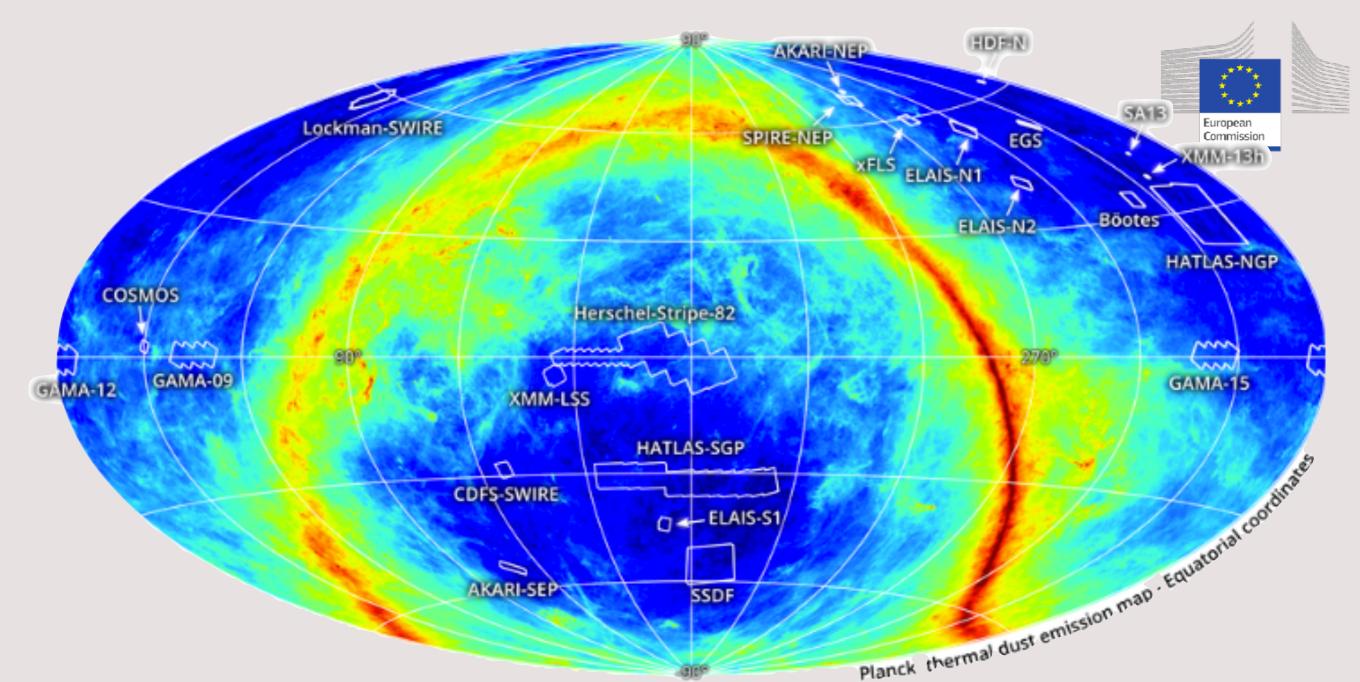


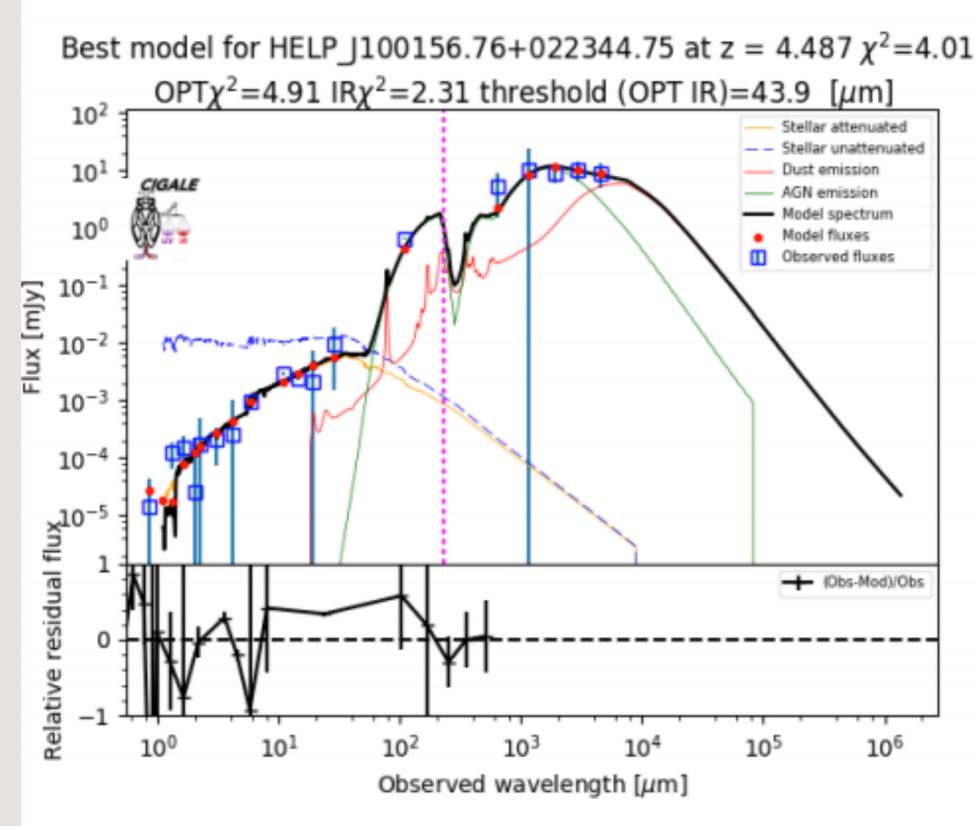


# HELP fields

herschel.sussex.ac.uk

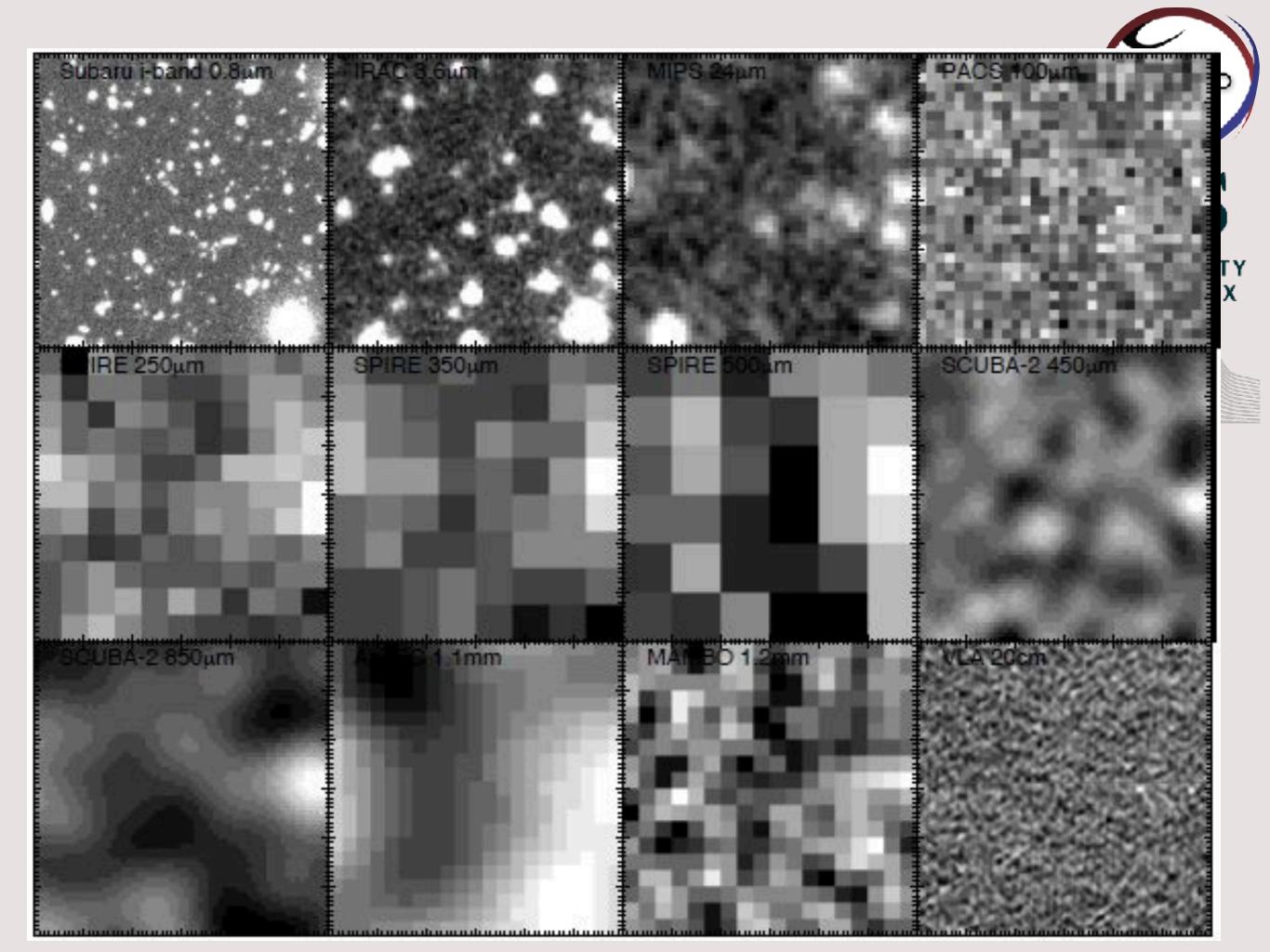








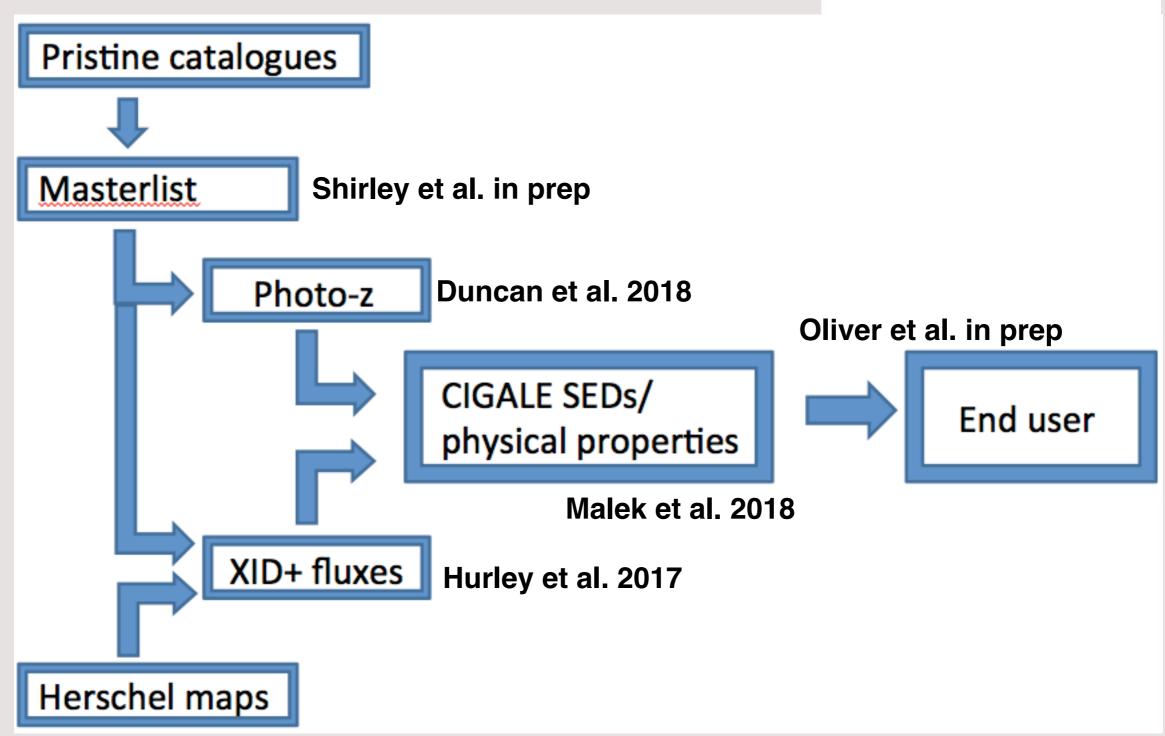






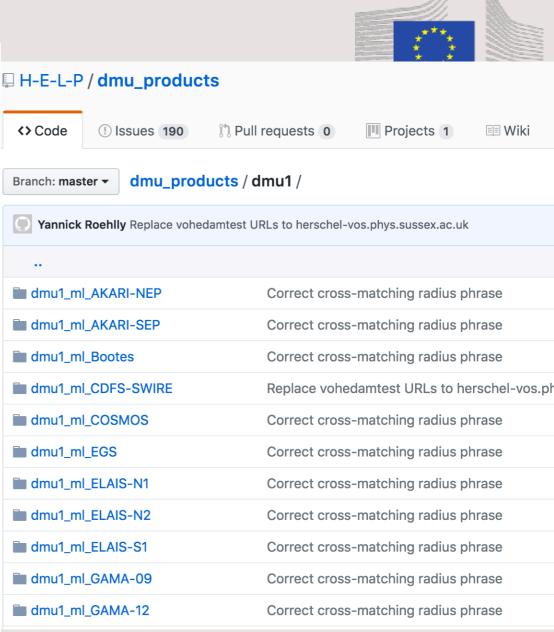






- herschel.sussex.ac.uk
- 60 Surveys
- 500 Gb catalogues
- github.com/H-E-L-P/dmu\_products
- 171 570 436 Objects!
- hedam.lam.fr/HELP





DMU#	Responsibility	Link
0	Pristine catalogues	dmu0
1	Masterlist data	dmu1
2	Field definitions	dmu2
3	Morphologies (Shapes & Sizes) of Objects	dmu3
4	Bright Star Mask	dmu4
5	Known Star Flag	dmu5
6	Optical photometry validation	dmu6
7	Optical photometry	dmu7
8	Radio data - LOFAR & FIRST/NVSS/TGSS	dmu8
9	Radio data - JVLA-DEEP & GMRT-DEEP	dmu9
10	Data Fusion	<u>dmu10</u>
11	Cross matching MIPS/PACS/SPIRE	<u>dmu11</u>

# Virtual Observatory at susseX (VOX)

- Allows complex queries to leverage the richness of the data.
- VO standards

https://herschel-vos.phys.sussex.ac.uk/



#### The Virtual Observatory at susseX (VOX)

Welcome to the Virtual Observatory at susseX (VOX). This is currently serving data from the Herschel Extragalactic Legacy Project (HELP) but will be extended to a range of astronomical datasets produced at Sussex and collaborating institutions. You can find out more about HELP here.

VOX is often the quickest way to access HELP data. However you can also browse the raw data <a href="here">here</a>. This includes large table files and images including some that is not served on VOX. You can read about the structure of the database on <a href="here">GitHub</a>.

In addition to the services listed below, on this site you can access <u>numerous tables</u> using TAP or form-based ADQL.

Please check out our site help.

#### Services available here

By Title	By Subject				
H					
	main catalogue master catalogue.	IQ			
		ing spectra from CIGALE i	Q		
	spectroscopic re endium of spectroscop	dshift catalogue i Q			
		nd SPIRE 250µm extraction) terschel Extragalactic Legacy Project.	<b>i</b> [Q]		
		nd SPIRE 350µm extraction) terschel Extragalactic Legacy Project.	i Q		
		nd SPIRE 500µm extraction) terschel Extragalactic Legacy Project.	iQ		
		E fluxes at SPIRE 250µm pos terschel Extragalactic Legacy Project.	sitions)	Q	
	hel map cutouts maps from the Hersch	i Q nel Space Observatory.			
	hel maps i Q maps from the Hersch	nel Space Observatory.			

# VO server – data access

- Through the web interface
  - Programmatically, e.g. in Python with pyVO.
- Through VO protocols e.g. with Topcat.



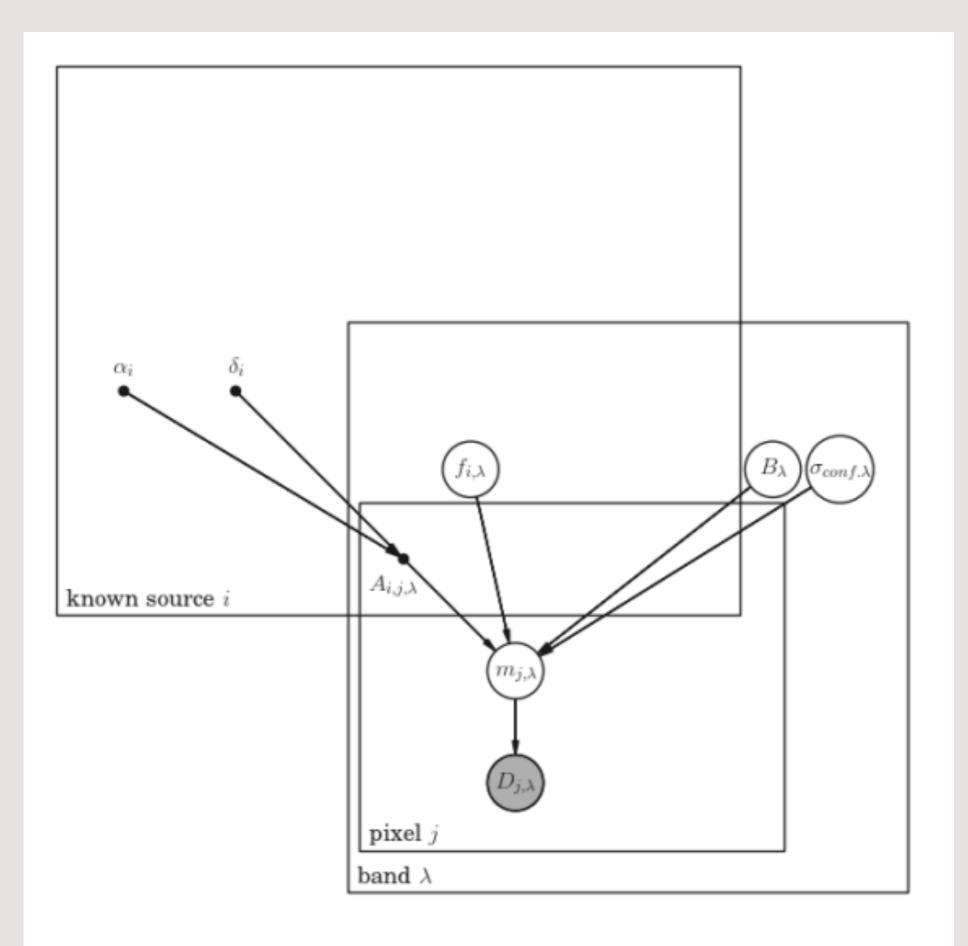






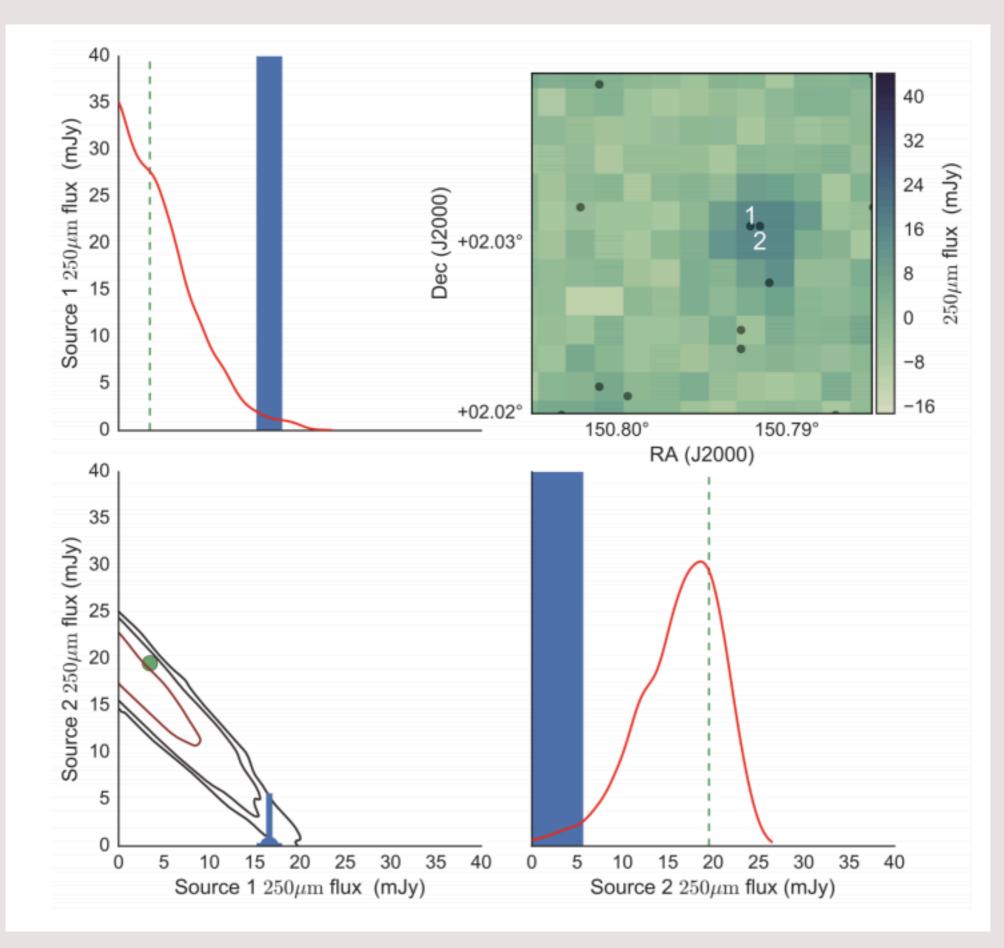
# XID+: The probabilistic de-blender for confusion dominated maps, Hurley et al. 2017

- Uses Bayesian Inference to get FULL posterior
- Provide a natural framework to introduce additional prior information
- Allows more accurate estimate of flux density errors for each source
- Provides a platform for doing science with the maps rather than catalogues



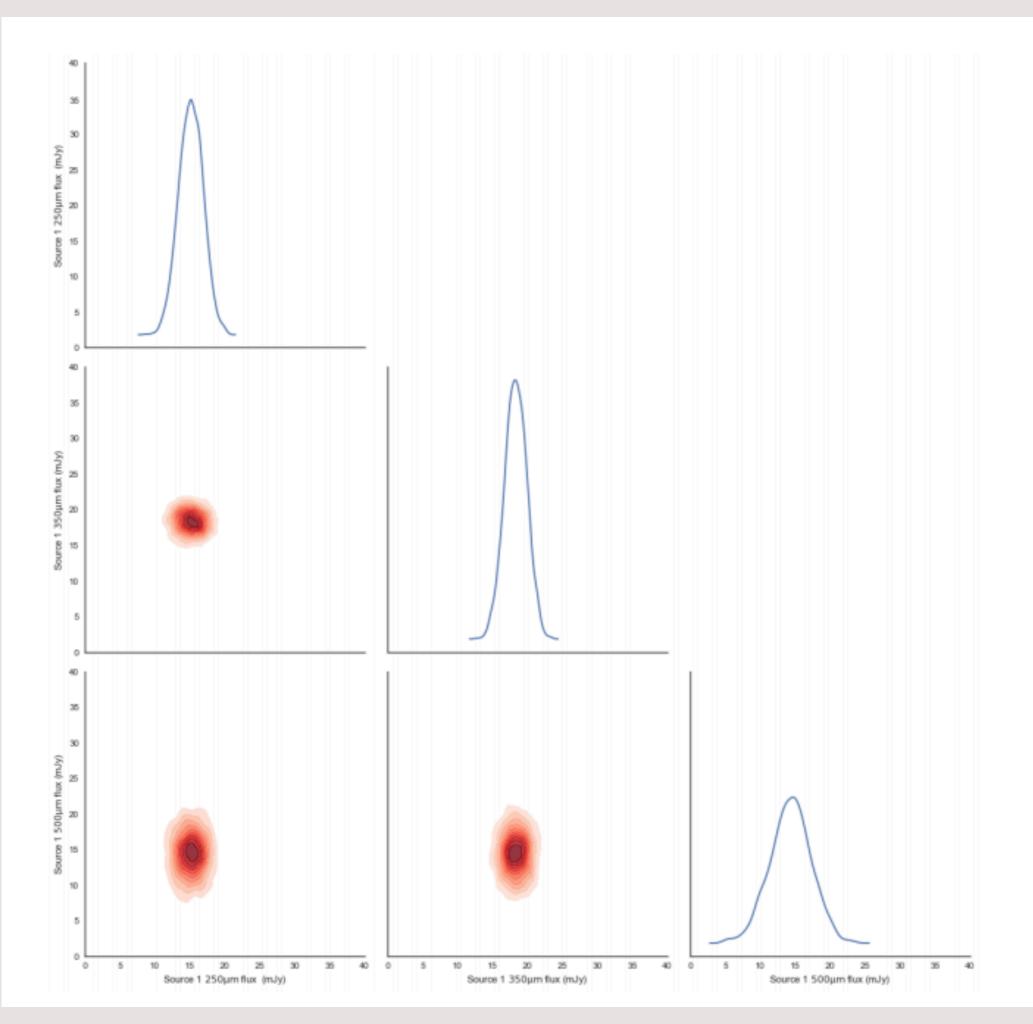






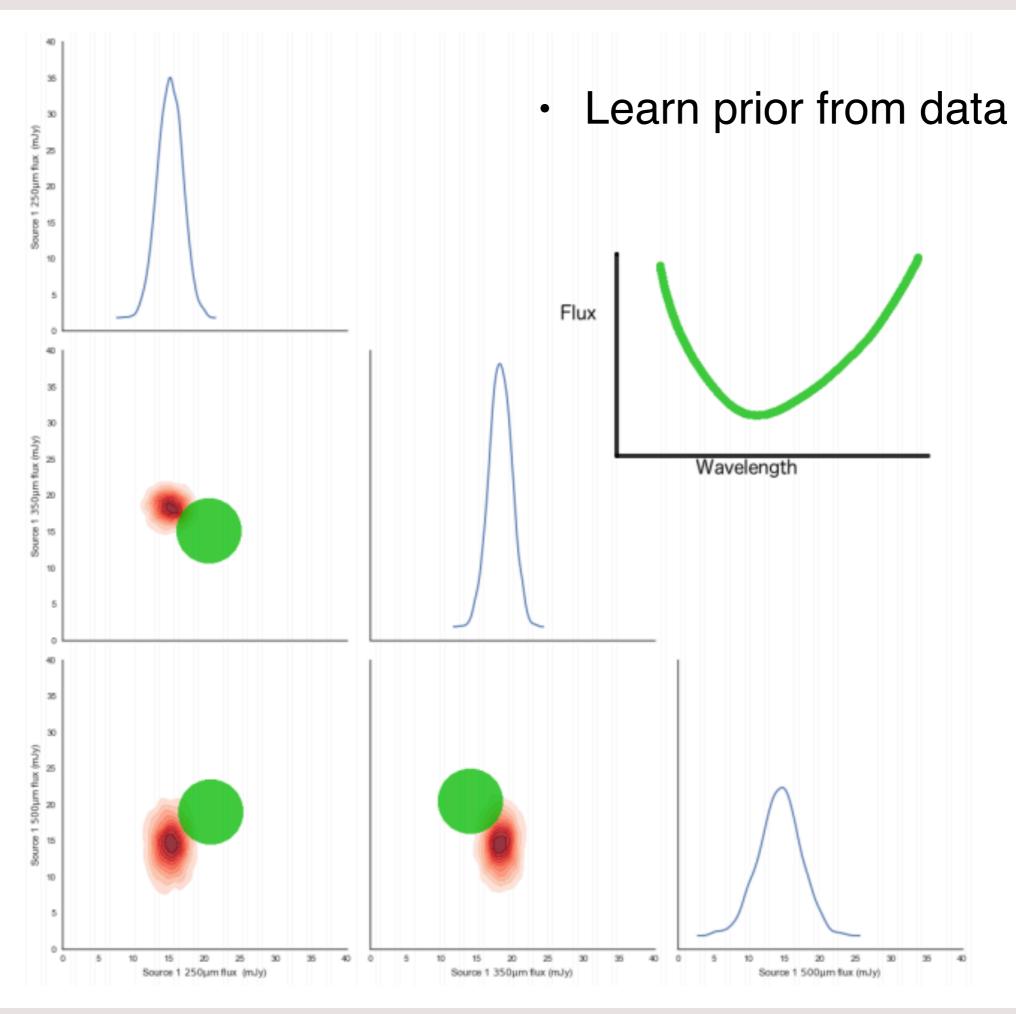








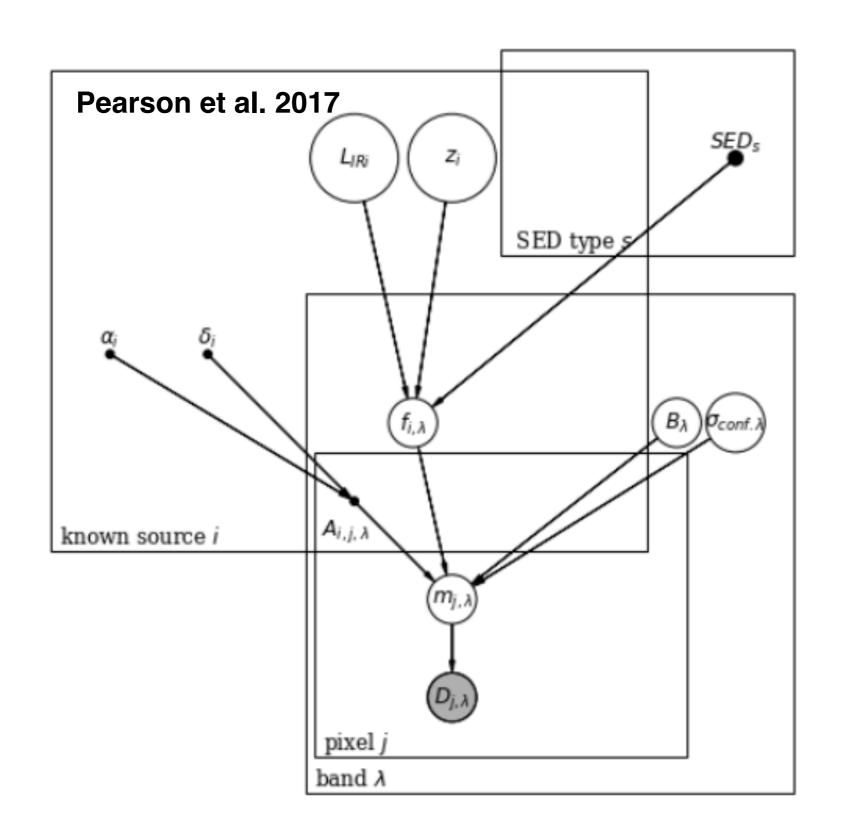








# XID+IR SED

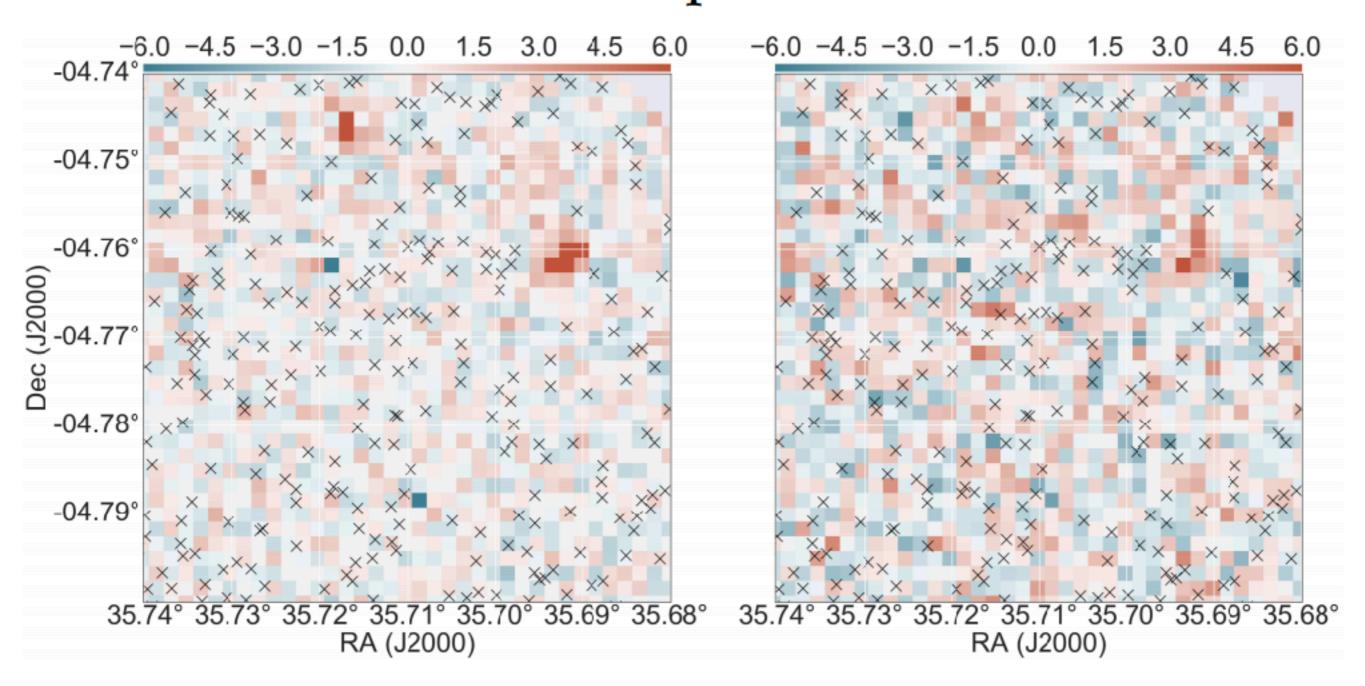








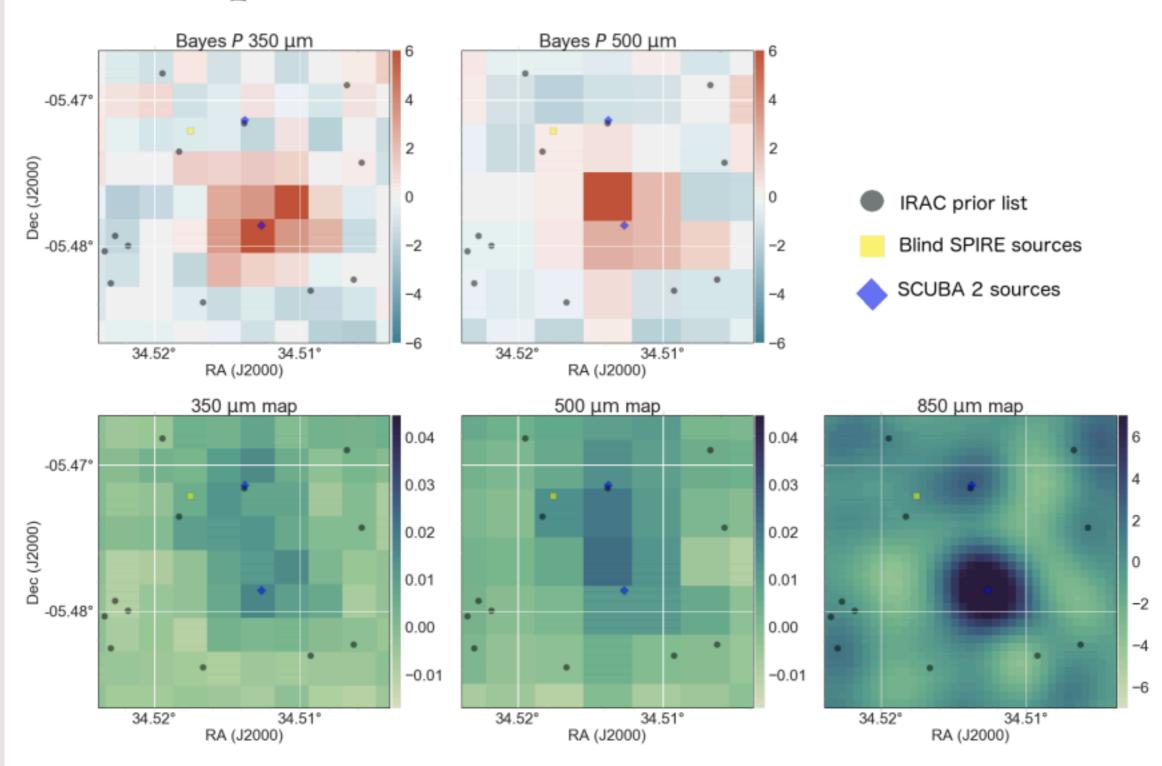
# XID+Bayesian P-value and standard MLM residual maps



# Maps reveal 'Hidden' FIR sources

SITY

SEX



Duncan et al. 2018

- photz -> Gaussian processes Almosallam et al. 2016
- 'A GP is a supervised non-linear regression algorithm'
- Needs a training set
- n, number of samples, with dimensionality d
- GPz models the distribution of functions that map those inputs on to the desired output
- training (0.8), validation (0.1), and test (0.1) samples





$$\mathbf{X} = \{x_i\}_{i=1}^n \in \mathbb{R}^{n \times d}$$

$$\mathbf{y} = \{y_i\}_{i=1}^n \in \mathbb{R}^n$$

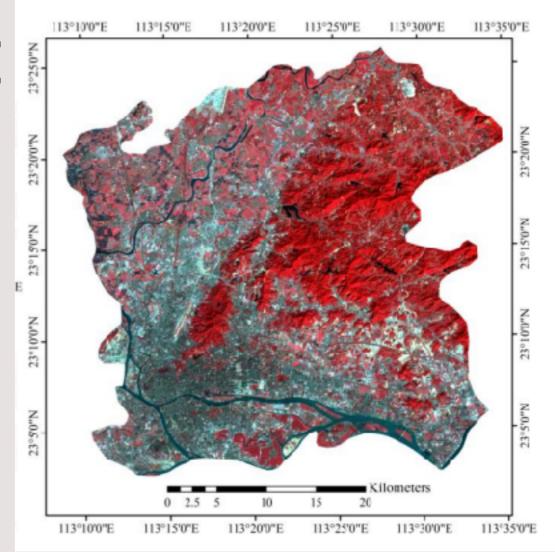
$$y_i = f(x_i) + \epsilon_i$$

Best model for HELP\_J100156.76+022344.75 at  $z = 4.487 \chi^2 = 4.01 E L P$  $OPT\chi^2 = 4.91 \text{ IR}\chi^2 = 2.31 \text{ threshold (OPT IR)} = 43.9 [\mu m]$ 10<sup>2</sup> Stellar attenuated Stellar unattenuated 10<sup>1</sup> Dust emission CIGALE AGN emission Model spectrum **ERSITY** 10<sup>0</sup> Model fluxes USSEX Observed fluxes  $10^{-1}$ Flux [m]y]  $10^{-2}$  $10^{-3}$  $10^{-4}$ Relative residual flux (Obs-Mod)/Obs 101 10<sup>2</sup>  $10^{3}$ 10<sup>5</sup> 10<sup>6</sup> 10° 10<sup>4</sup> Observed wavelength [µm]





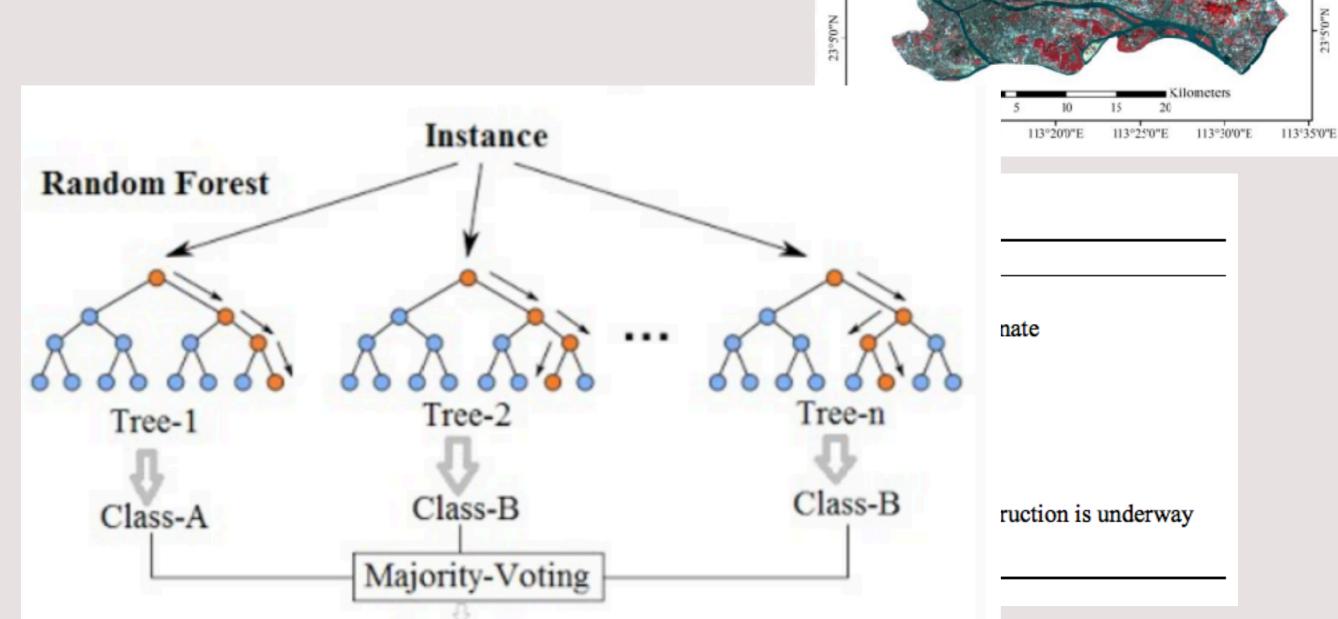
- training sample
- 500 pixels
- mean, max, sigma for every wavelength band



	Land-Use Ty	ypes Description
	Water	Water bodies such as reservoirs, ponds and river
	Residential a	rea Residential areas where driveways and roof tops dominate
Li et al.	2014 Natural for	est Large area of trees
Li Ct ai.	Orchard	Large area of fruit trees planted
	Farmland	Fields where vegetables or crops grow
	Industrial/comr	nercial Lands where roof tops of large buildings dominate
	Cleared land/Lar construction	Lands where vegetation is denuded or where the construction is underway
	Idle land	Lands where no vigorous vegetation grows
	idie land	Lands where no vigorous vegetation grows



- 85% accurate 4 bands
- 88% accurate 6 bands



113°15'0"E

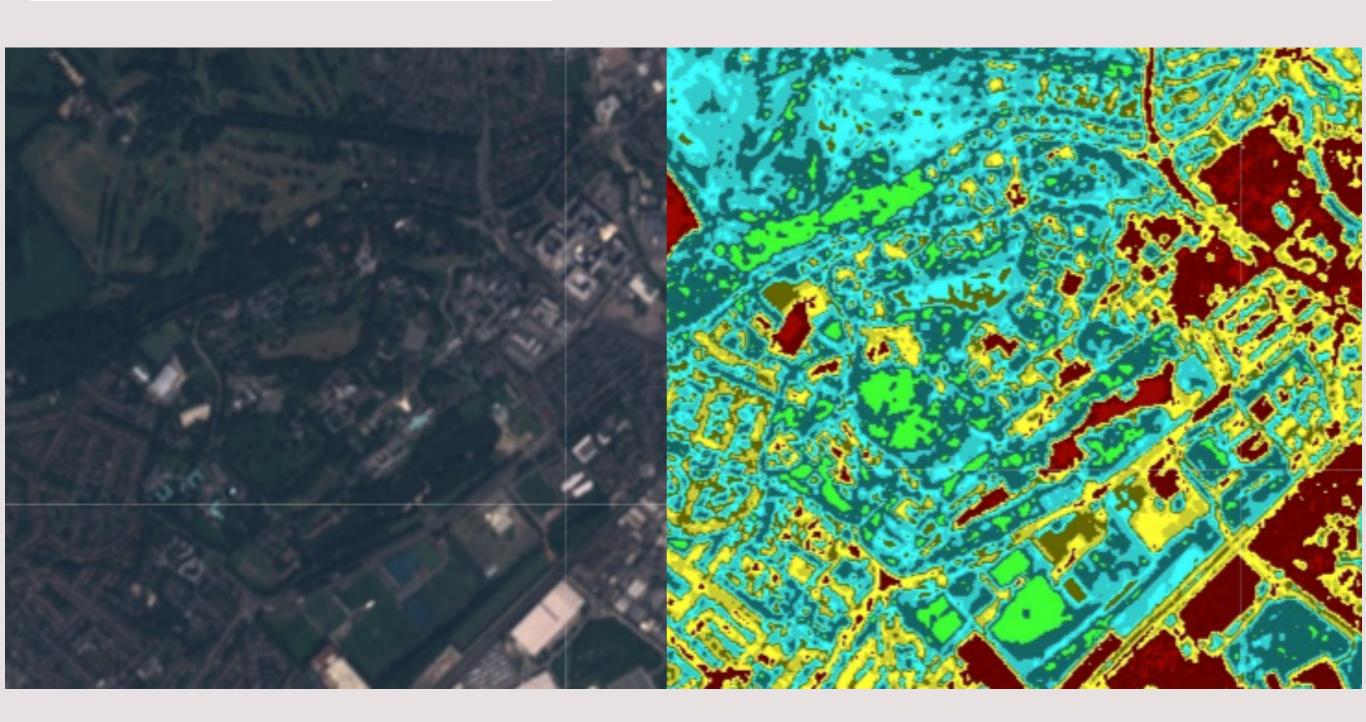
23°150"N

23\*10\*0\*N

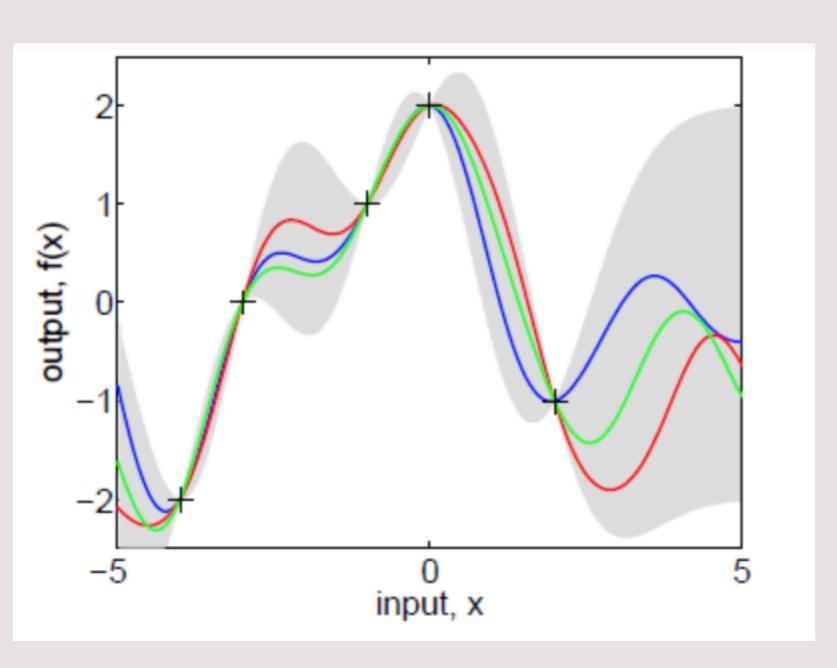
113°20'0"E

113°25'0"E







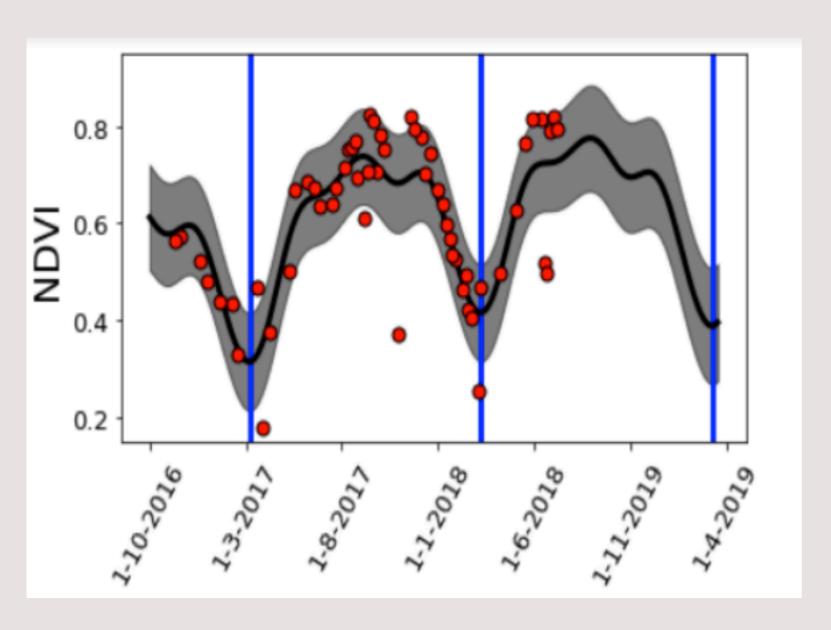


- GP
- scikit learn
- RBF and periodic kernel



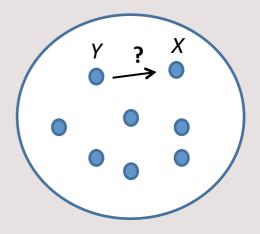
```
from sklearn.gaussian_process import GaussianProcessRegressor
from sklearn.gaussian_process.kernels import RBF, WhiteKernel, ExpSineSquared
```





- GP
- scikit learn
- RBF and periodic kernel
- Include spatial information?





#### Granger causality

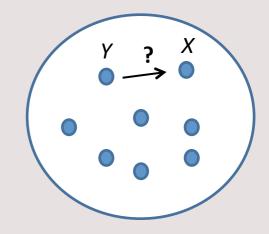
- **-Causality based on prediction**: Does past of Y help predict future of X?
- -Over and above the past of X itself, and the past of other variables?
- -Prediction by regression, using past observations as predictors.
- -Compare prediction errors (residuals) for regressions including and excluding Y

$$X_1(t) = \sum_{j=1}^p A_{11,j} X_1(t-j) + \sum_{j=1}^p A_{12,j} X_2(t-j) + E_1(t)$$

$$X_2(t) = \sum_{j=1}^p A_{21,j} X_1(t-j) + \sum_{j=1}^p A_{22,j} X_2(t-j) + E_2(t)$$



### **Granger** causality



- -Apply to NDVI / VCI vegetation state index time-series.
- -Subtract off seasonal fluctuations and overall linear trend.
- -Map dependence of GC on distance between predictor and predictee pixels.
- -Compare performance of linear regression models with non-linear regression models such as random forests.
- -Compare prediction of Principal Component time-series with prediction of regional mean time-series.



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github.com/H-E-L-P/dmu\_products

hedam.lam.fr/HELP

